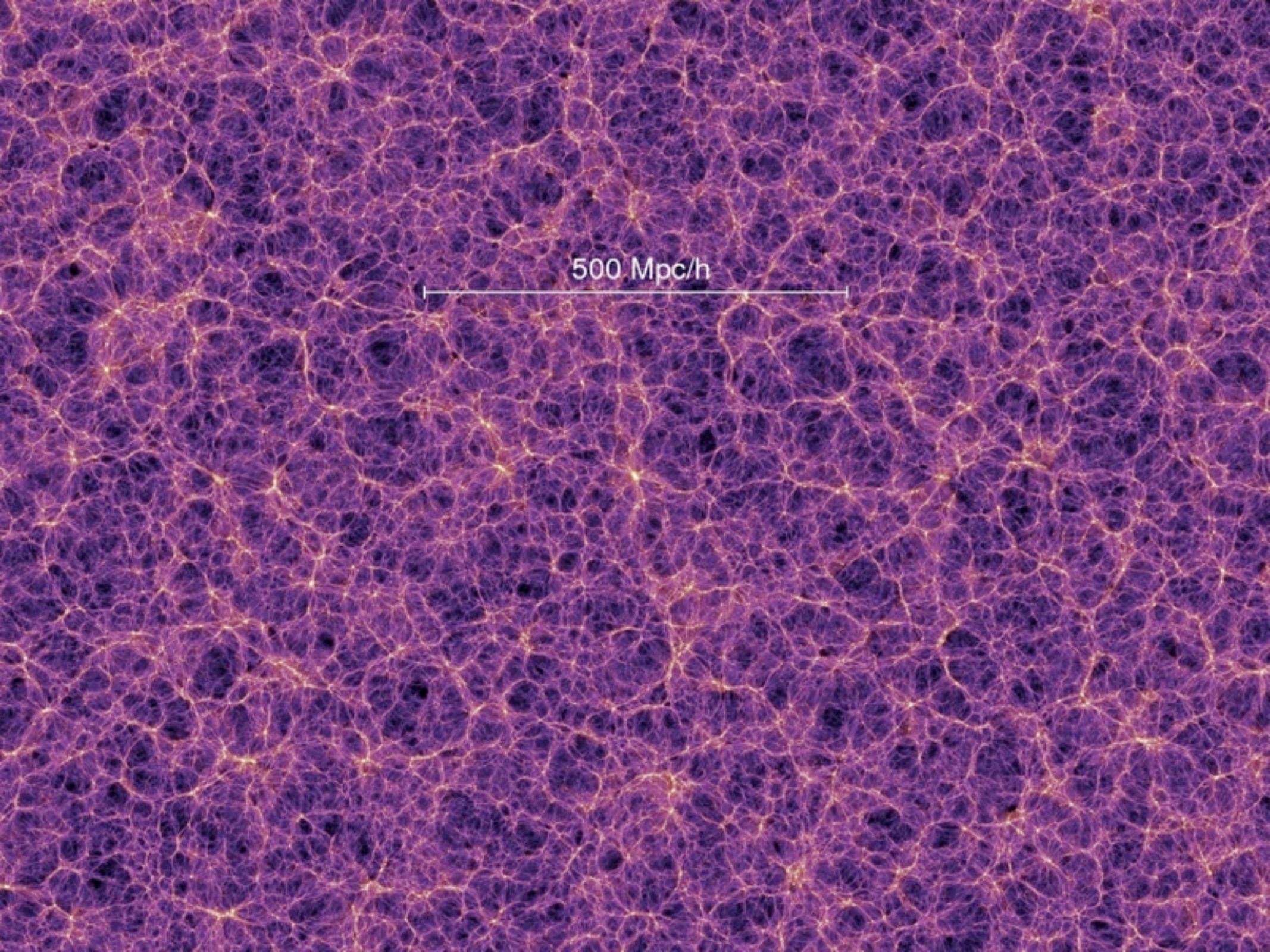
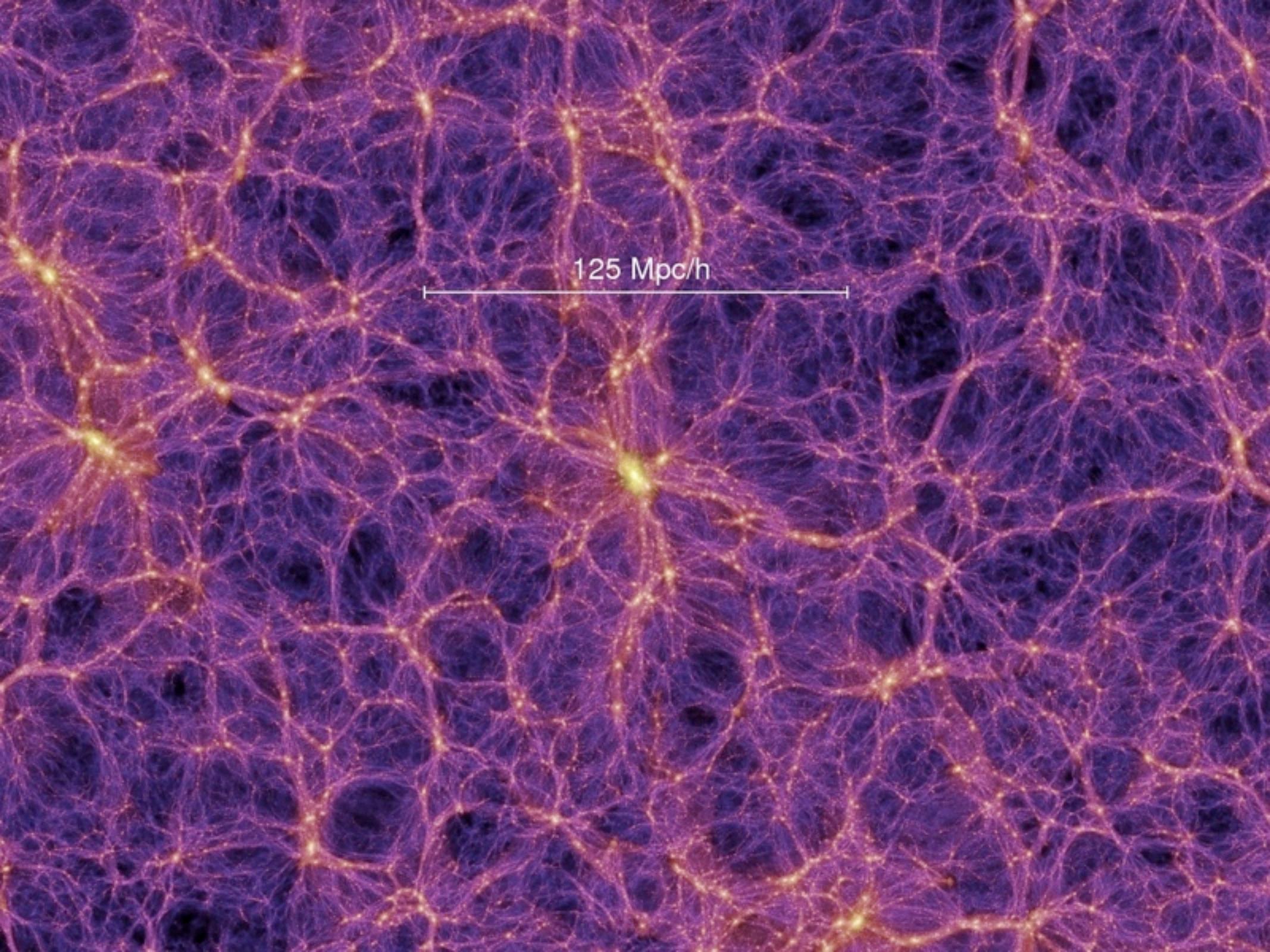
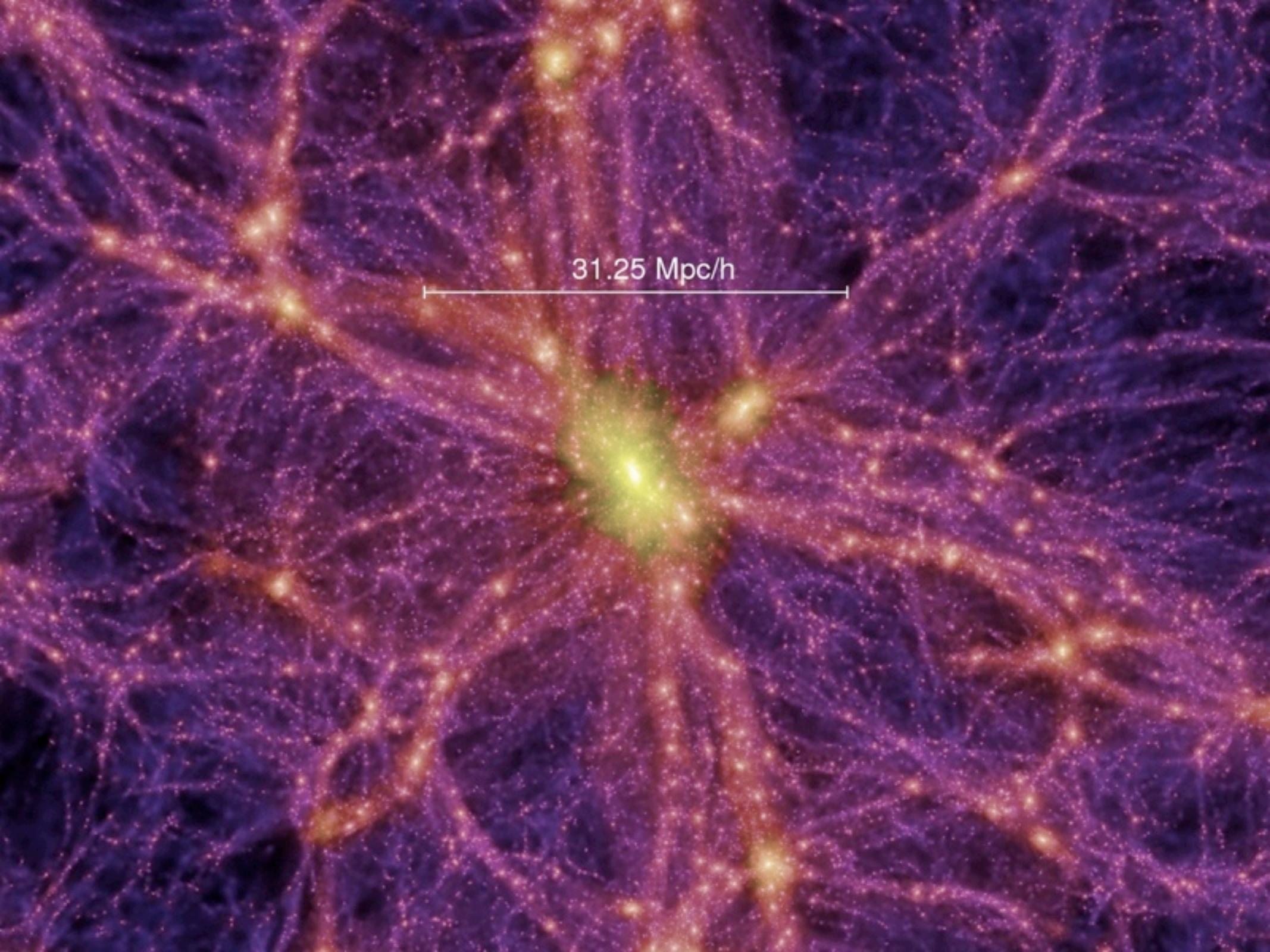


distribution of galaxies in space



500 Mpc/h

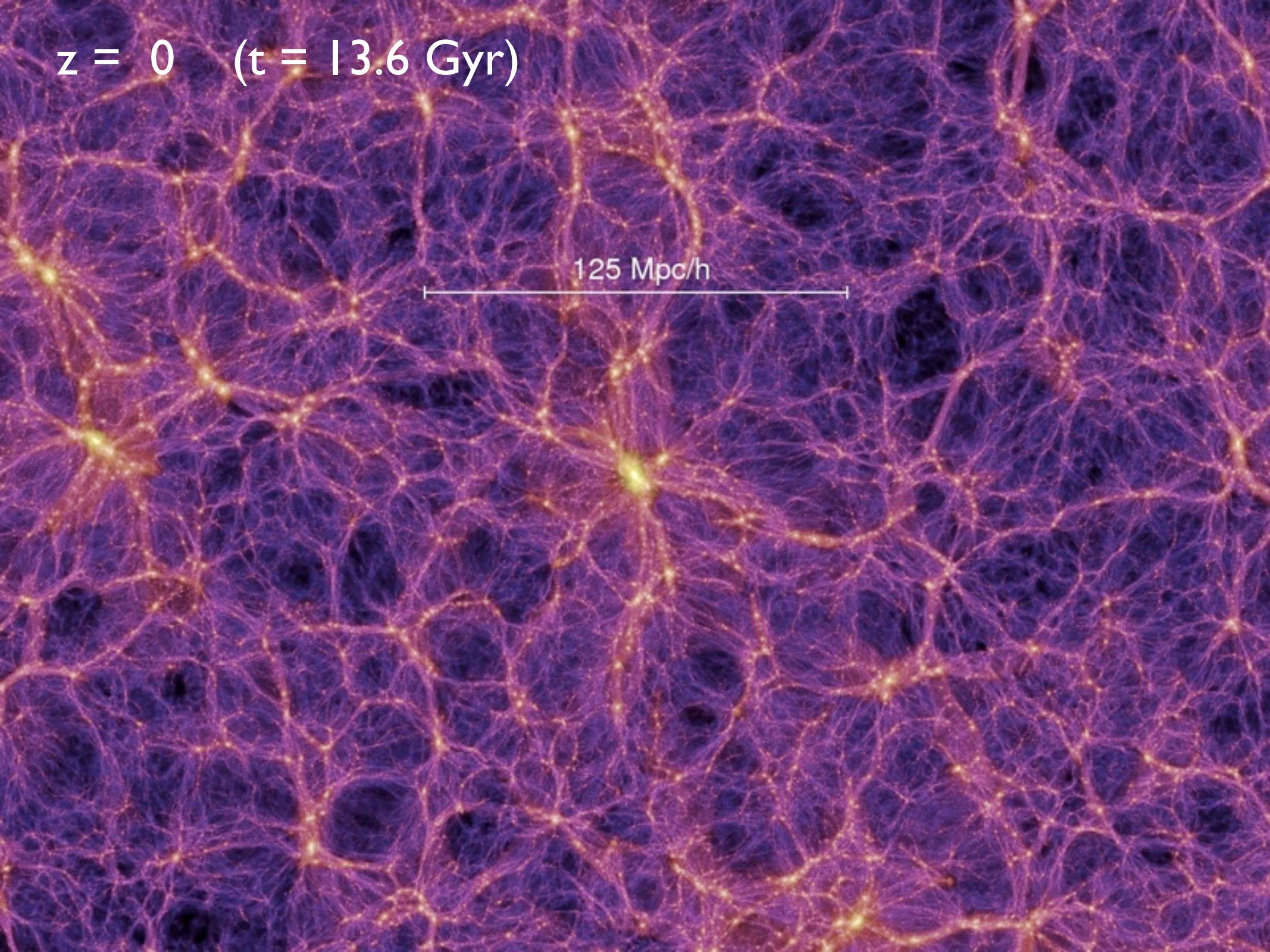




A wide-field image of a galaxy cluster, likely Abell 1758, showing a dense concentration of galaxies. The galaxies are color-coded by their temperature or velocity, ranging from red (lower temperature/velocity) to blue (higher temperature/velocity). A prominent green and yellow central region indicates the core of the cluster. A scale bar in the upper right corner indicates a distance of 31.25 Mpc/h.

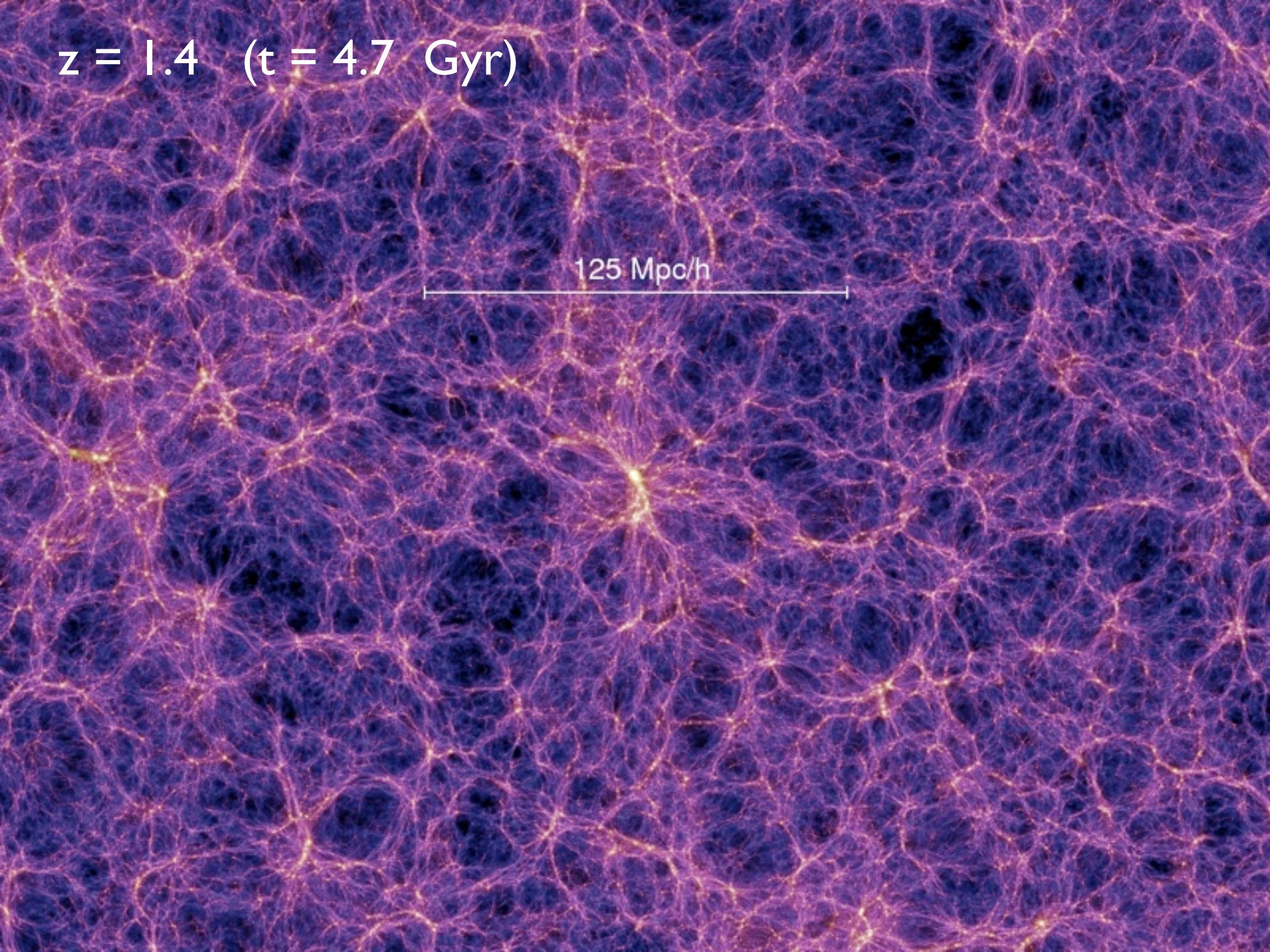
31.25 Mpc/h

$z = 0$ ($t = 13.6$ Gyr)



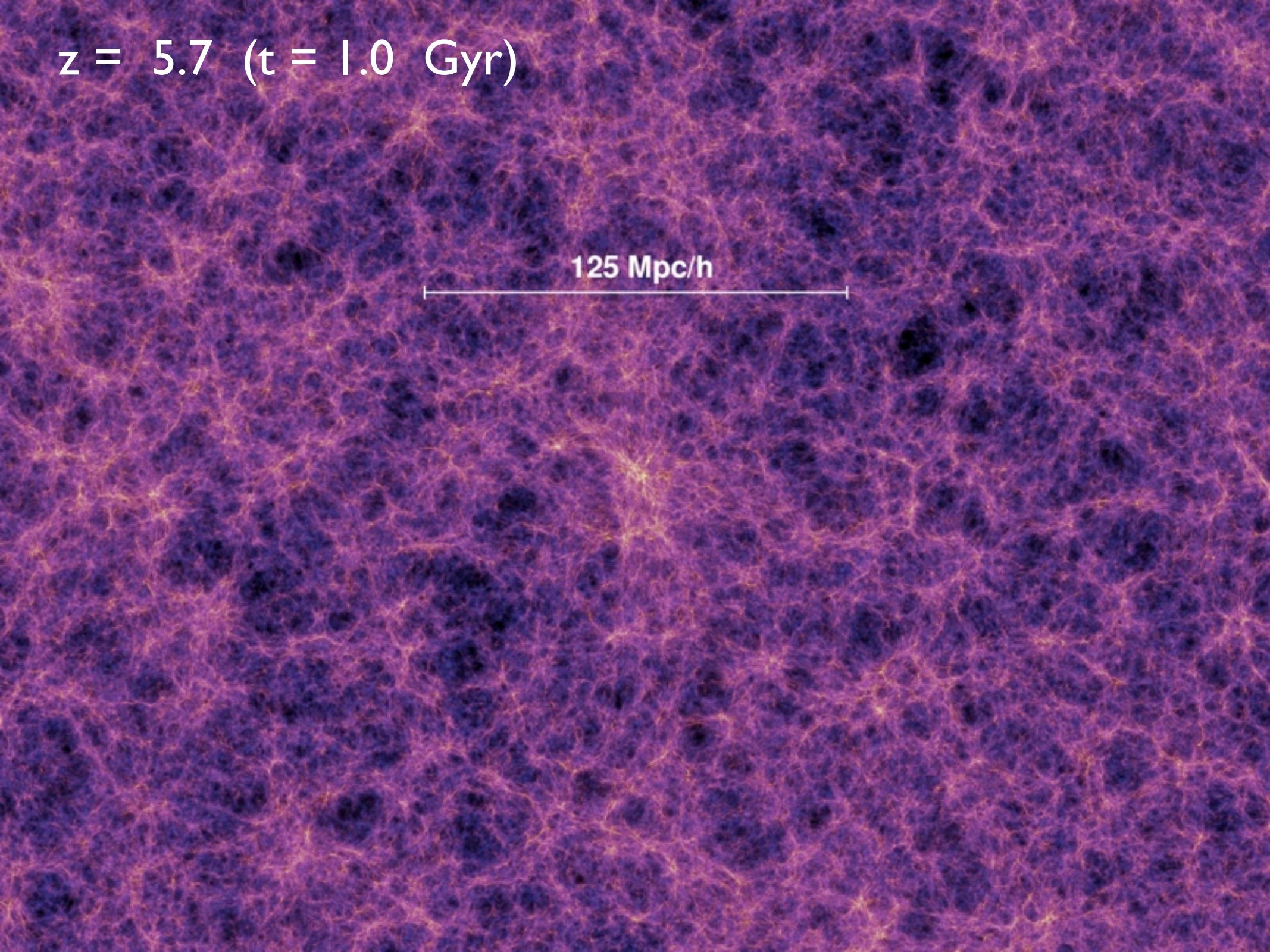
125 Mpc/h

$z = 1.4$ ($t = 4.7$ Gyr)



125 Mpc/h

$z = 5.7$ ($t = 1.0$ Gyr)



125 Mpc/h

$z = 18.3$ ($t = 0.21$ Gyr)



125 Mpc/h

Sun:
0.5 degree diameter

R = 696,000 km
d = 149,000,000 km

Royal Observatory of Belgium - SIDC

DATE-OBS= '19/06/2007'

TIME = '10:42:33'

INSTRUME= "WHITE-LIGHT"

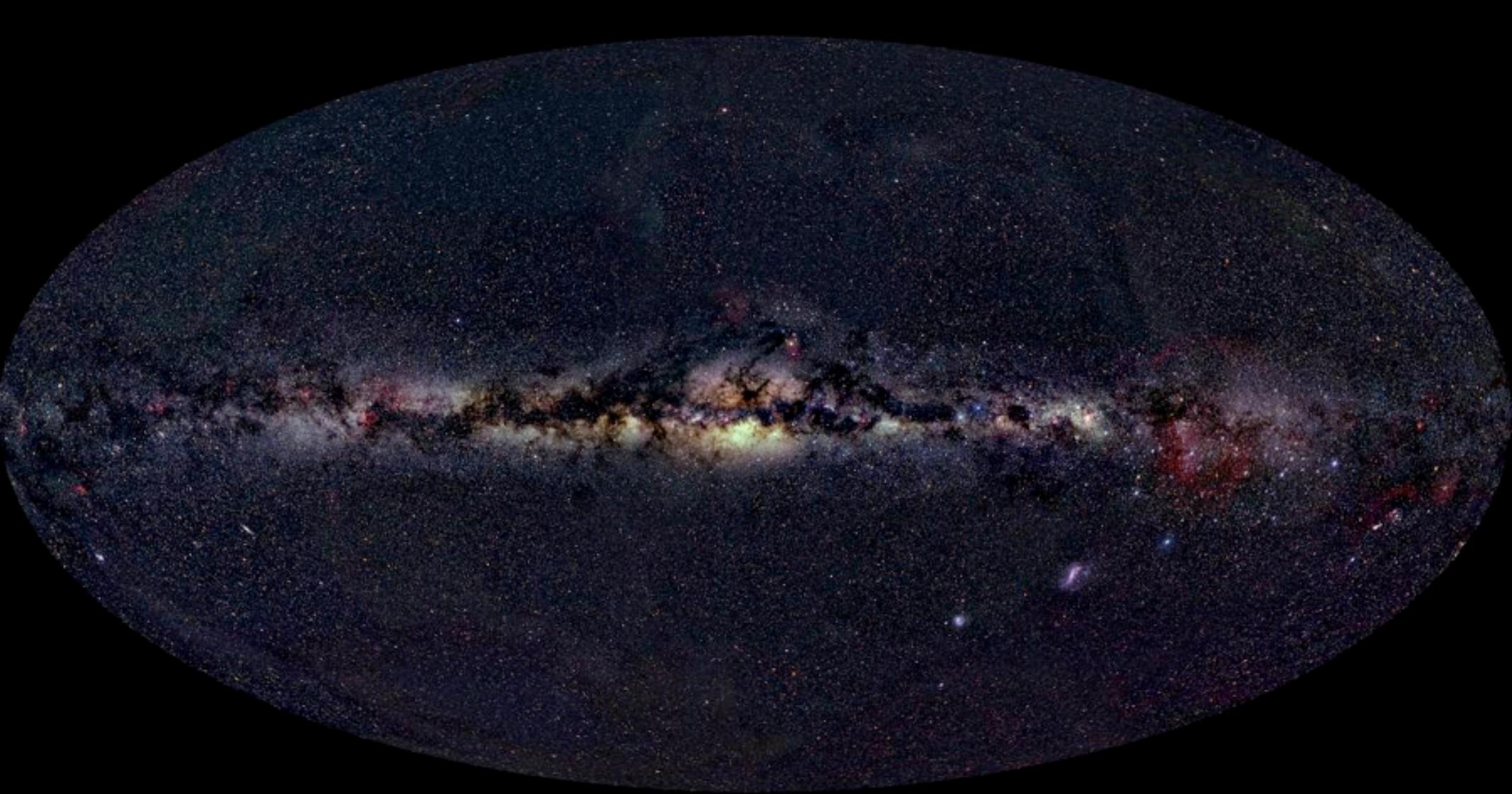
EXPTIME = 0.061000

α Centauri

28 deg

$d = 1.35$ parsec

Alpha Centauri and the Southern Cross



© 2000, Axel Mellinger

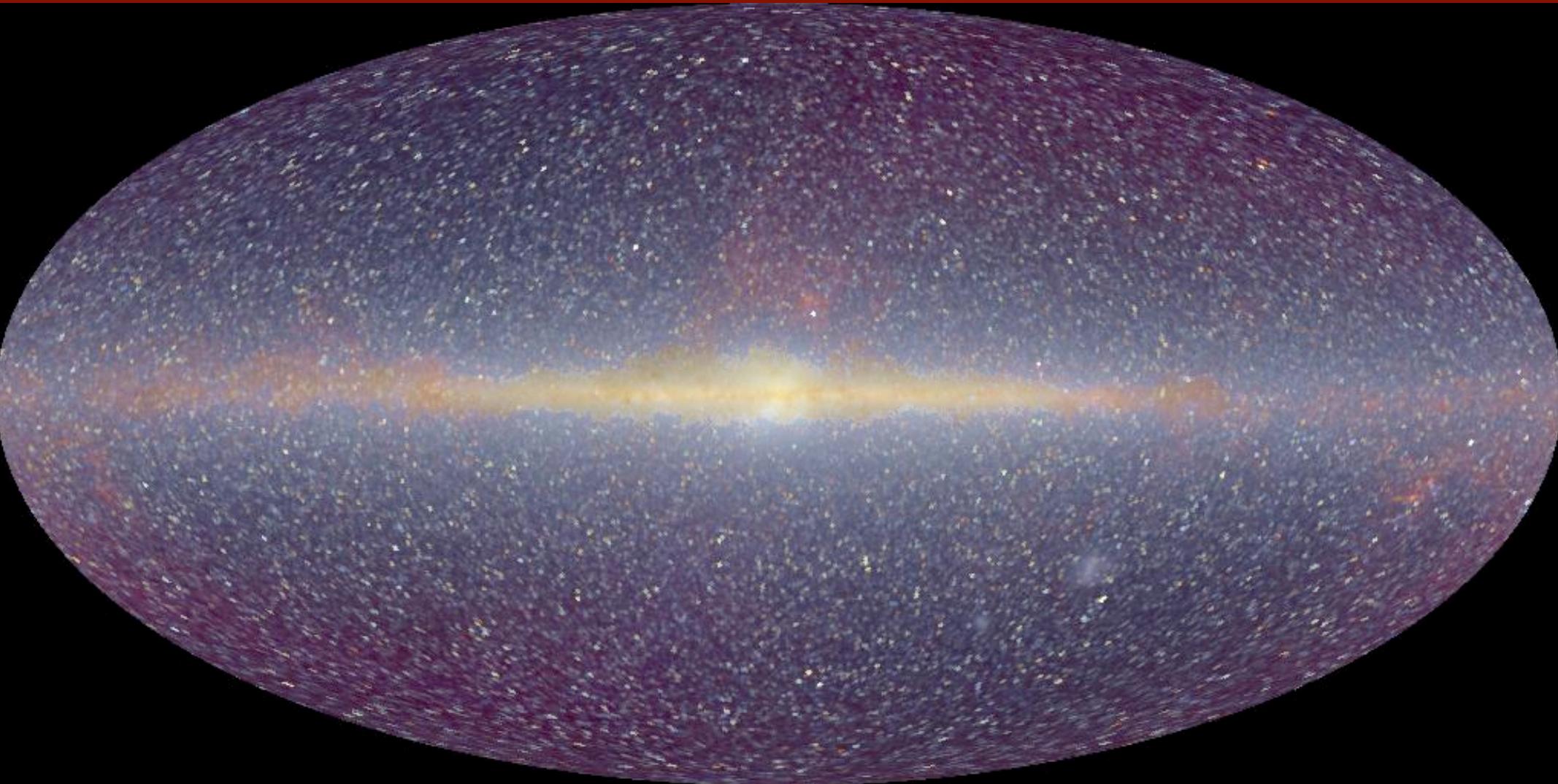
Cygnus

Sagittarius

Crux

Orion

all-sky map: Milky Way disk is aligned along the equator

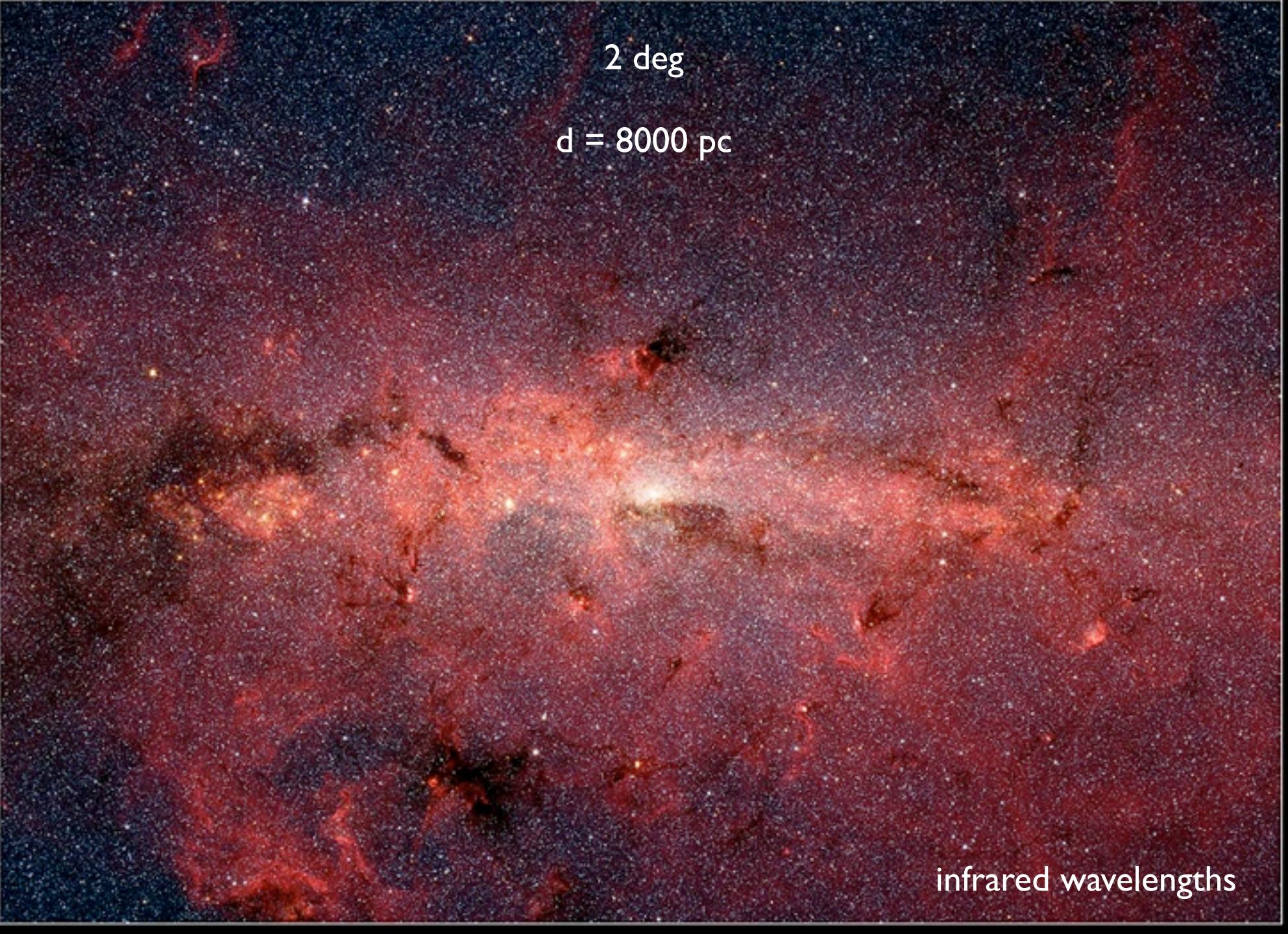


Diffuse Infrared Background Explorer
view of the Milky Way

Interstellar dust can be opaque at visible wavelengths,
but transparent at infrared wavelengths

Galactic center in Sagittarius: 65 deg





2 deg

$d = 8000 \text{ pc}$

infrared wavelengths

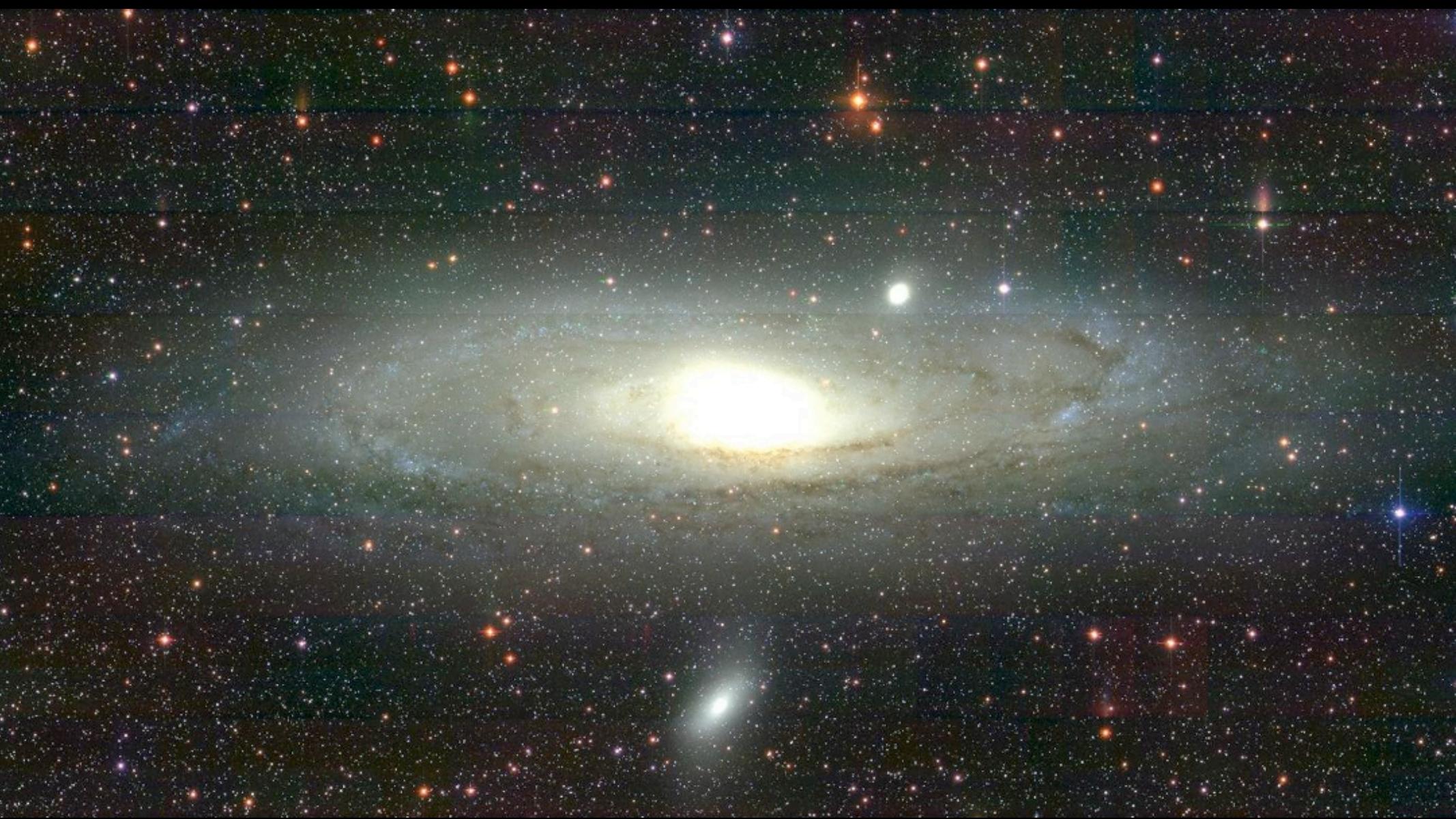
The Center of the Milky Way Galaxy

NASA / JPL-Caltech / S. Stolovy (Spitzer Science Center/Caltech)

Spitzer Space Telescope • IRAC

ssc2006-02a

M31 = Andromeda: 3 deg
d = 780 kpc



1932: brightest galaxies
(Shapley-Ames catalog)

$d \sim 10$ Mpc, $N \sim 1000$

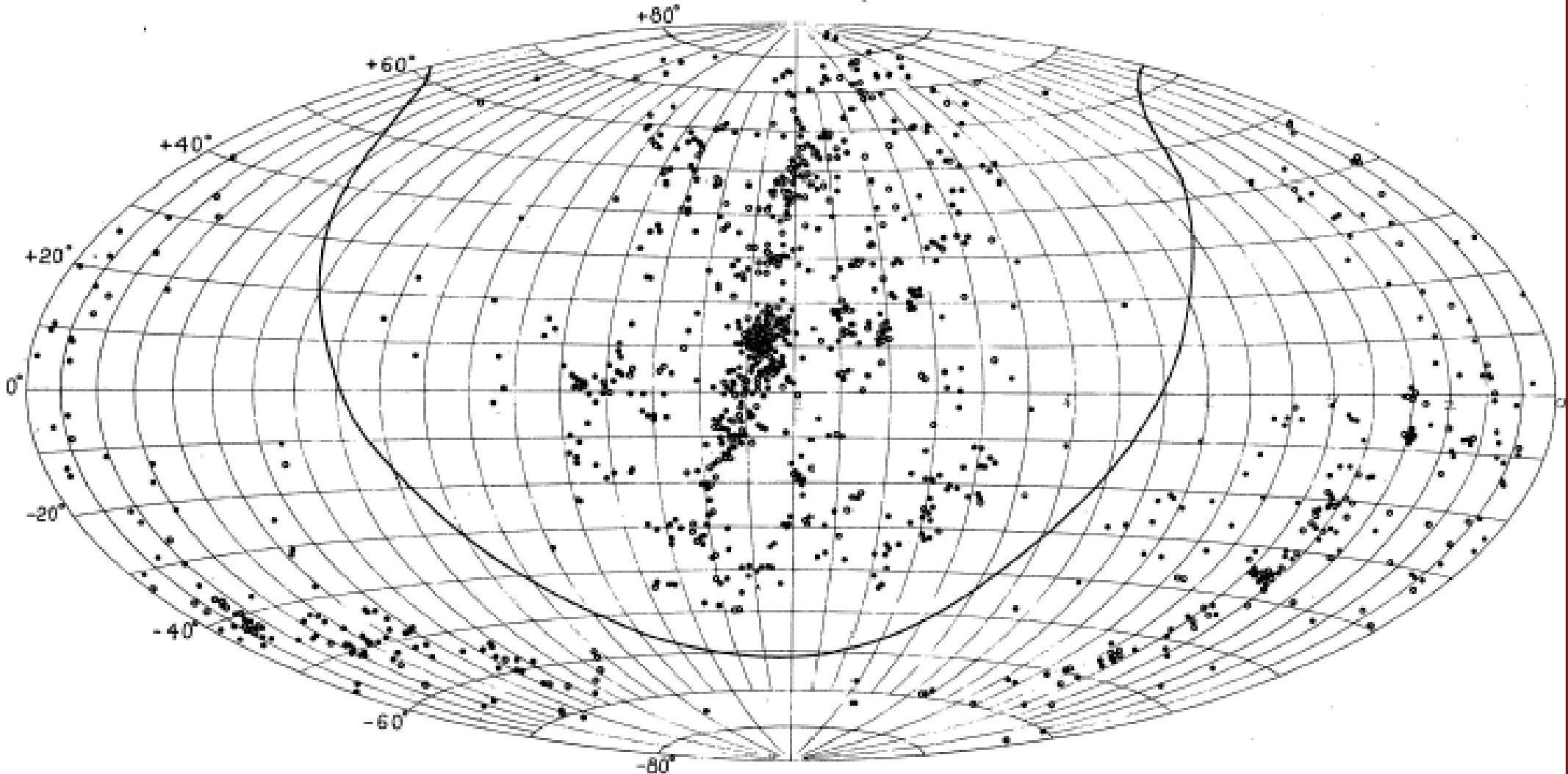
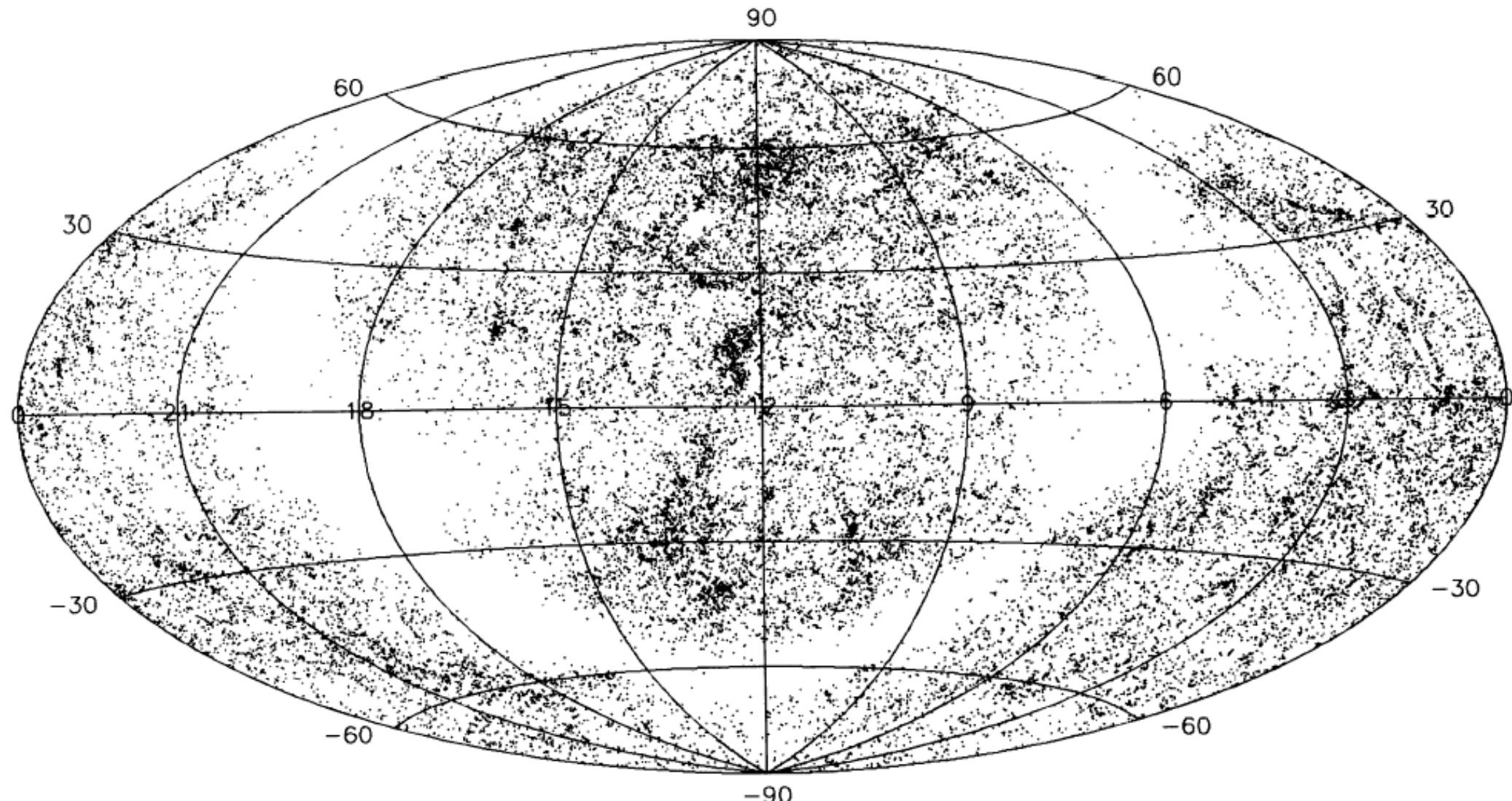


Figure 4

1970's: fainter galaxies
(Zwicky catalog)

$d \sim 30$ Mpc, $N \sim 27000$





28.8 deg

Coma Berenices and Virgo

Coma Berenices

28.8 deg

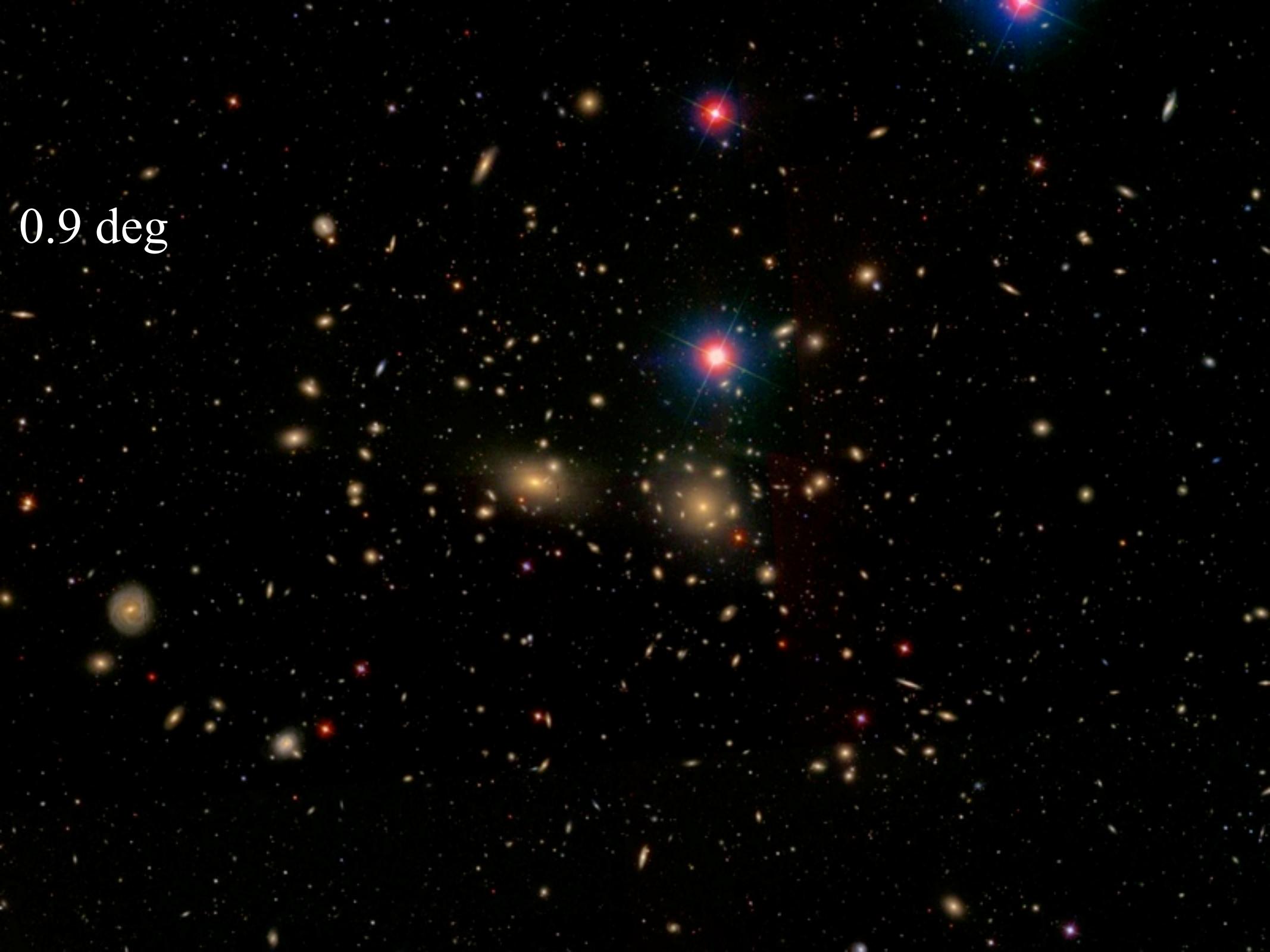
14.4 deg





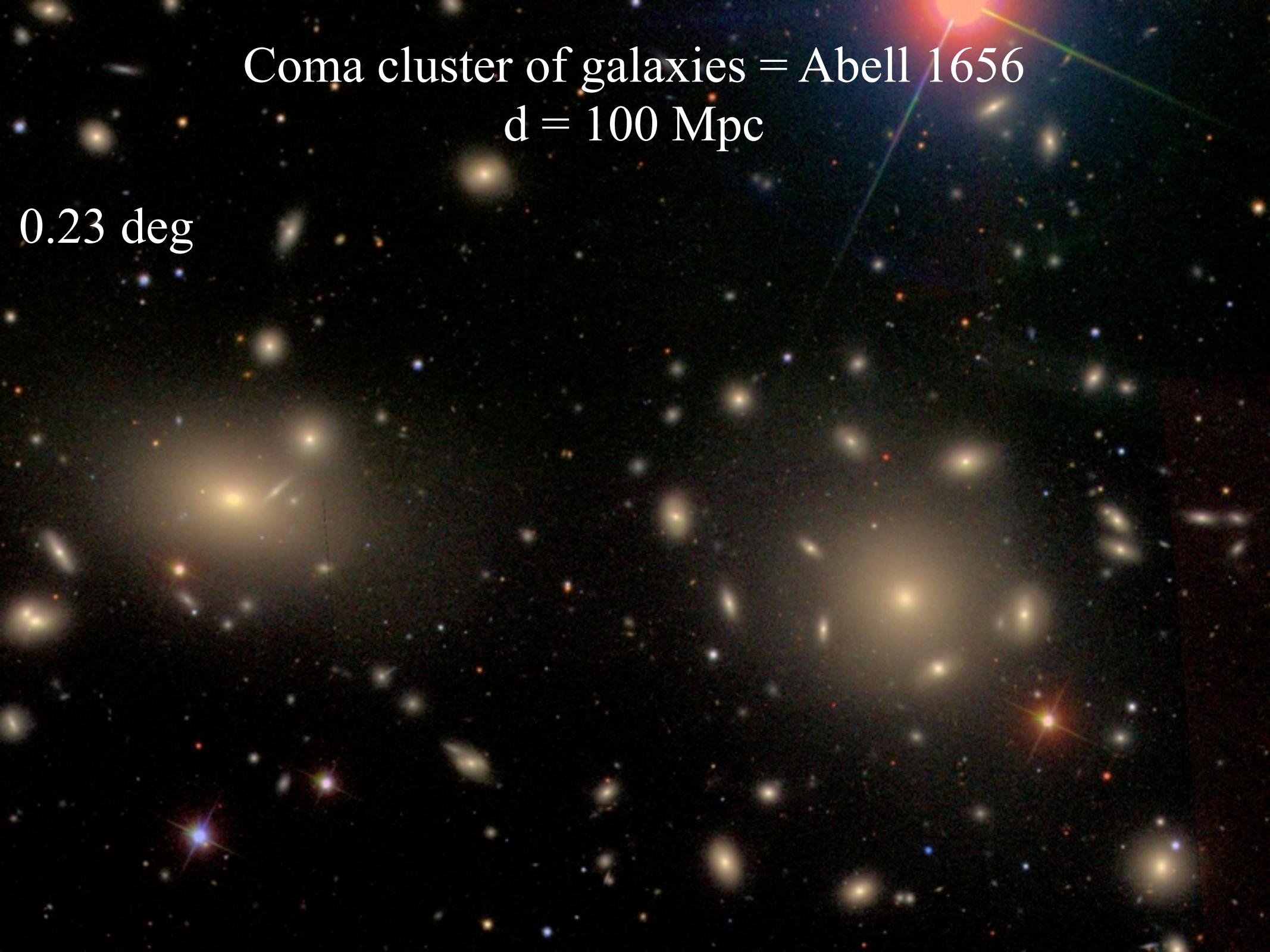
3.6 deg

0.9 deg

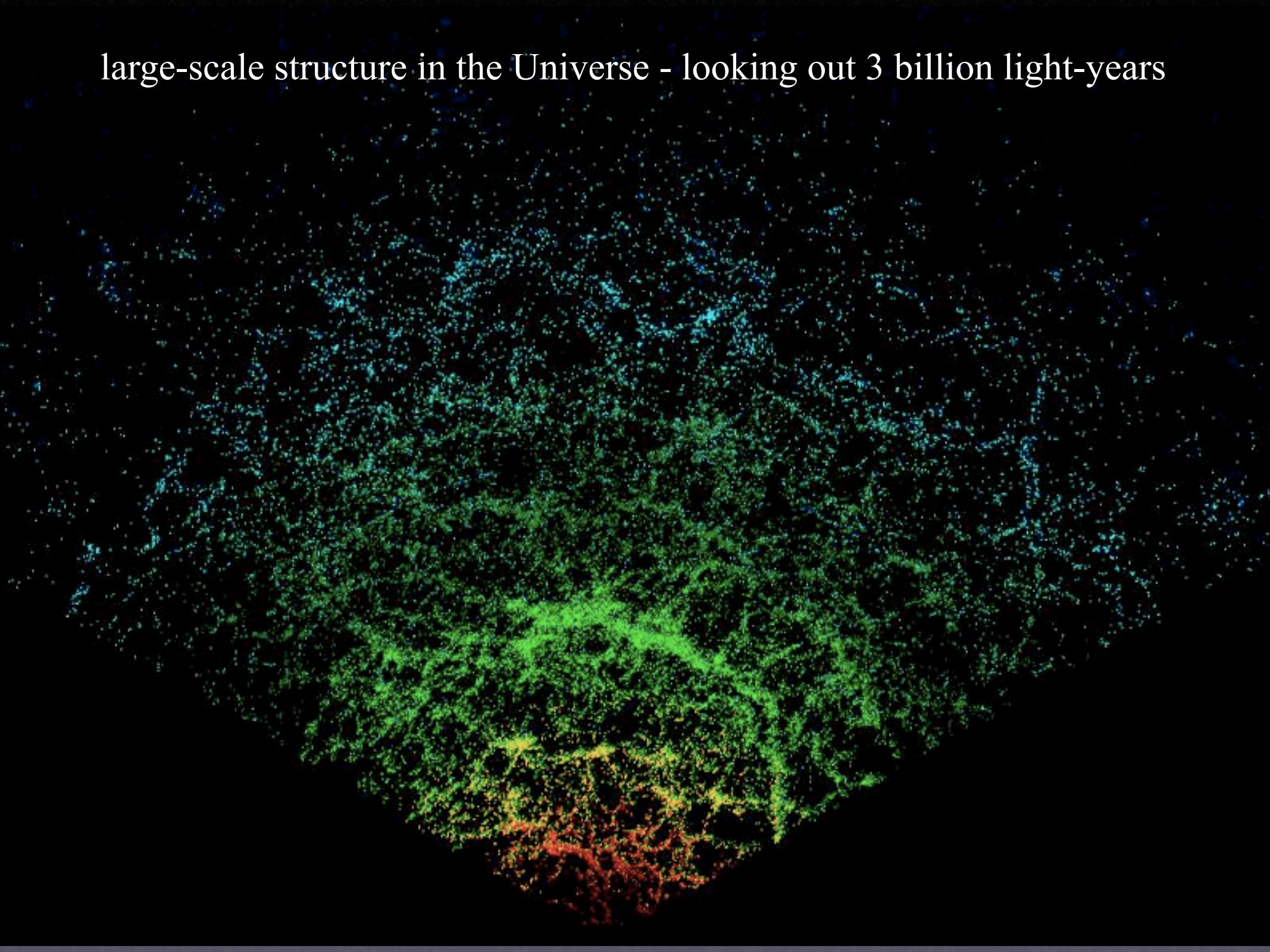


Coma cluster of galaxies = Abell 1656
 $d = 100 \text{ Mpc}$

0.23 deg



large-scale structure in the Universe - looking out 3 billion light-years



Sun:
0.5 degree diameter

R = 696,000 km
d = 149,000,000 km

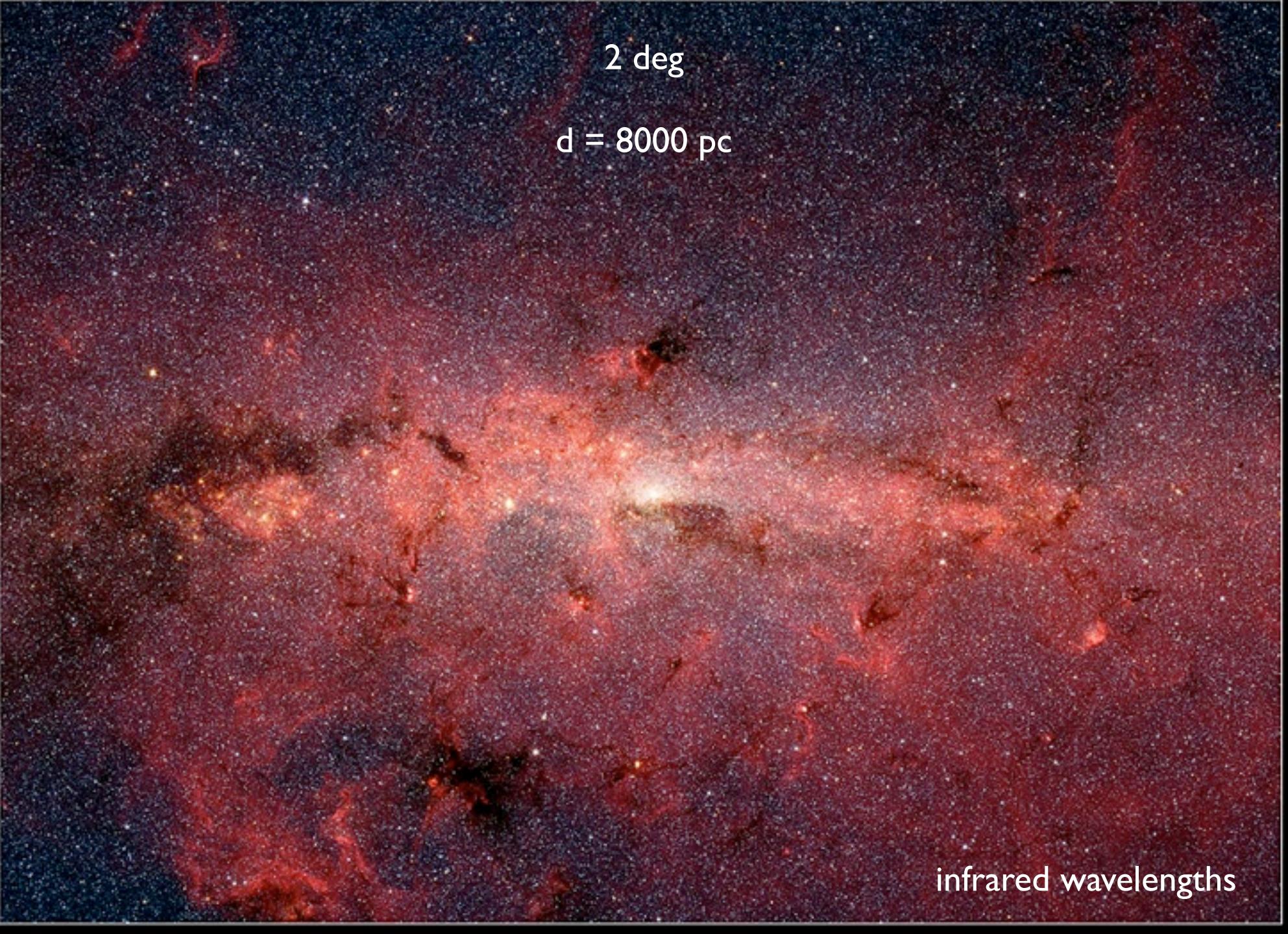
Royal Observatory of Belgium - SIDC

DATE-OBS= '19/06/2007'

TIME = '10:42:33'

INSTRUME= "WHITE-LIGHT"

EXPTIME = 0.061000



2 deg

$d = 8000 \text{ pc}$

infrared wavelengths

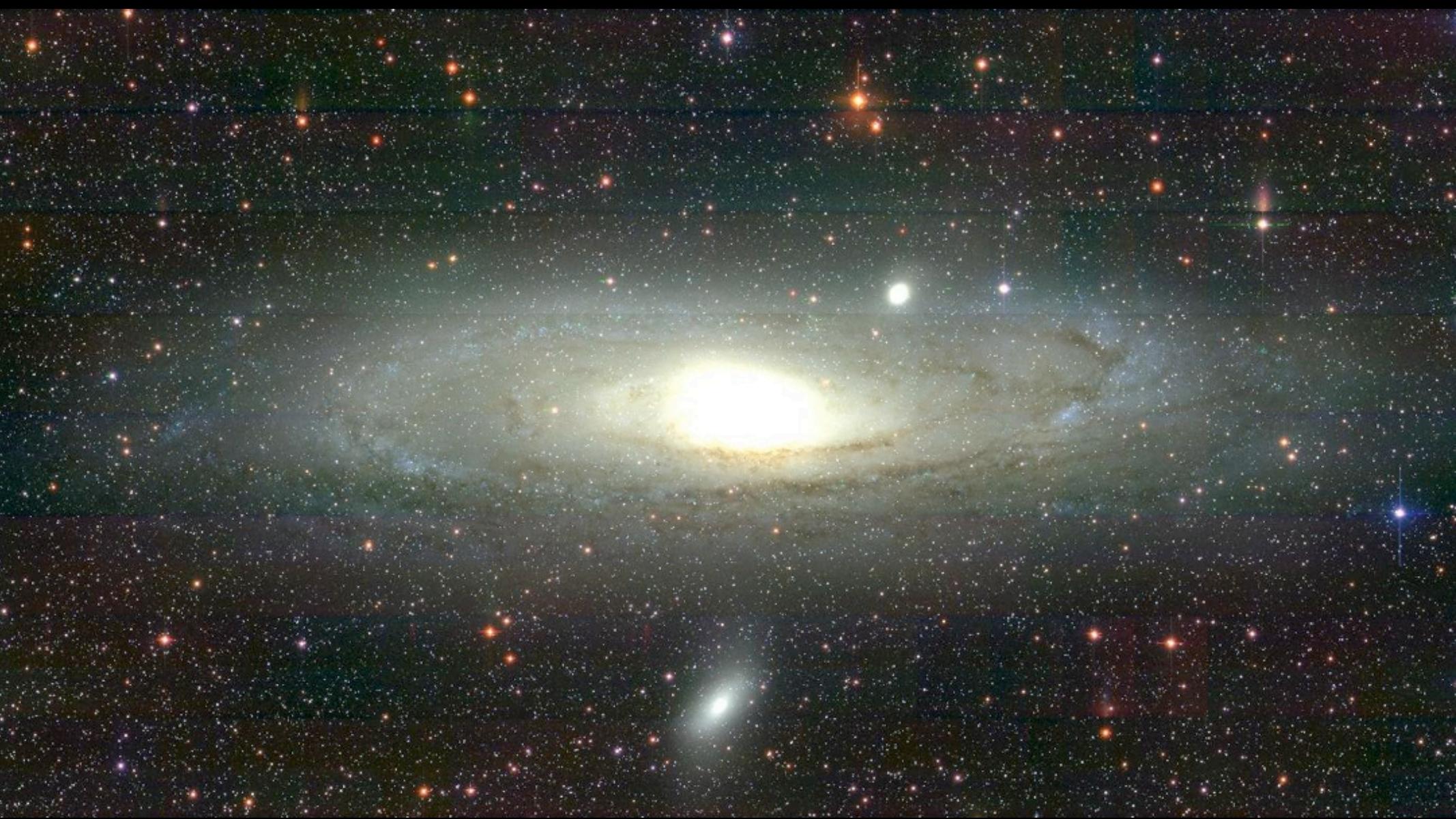
The Center of the Milky Way Galaxy

NASA / JPL-Caltech / S. Stolovy (Spitzer Science Center/Caltech)

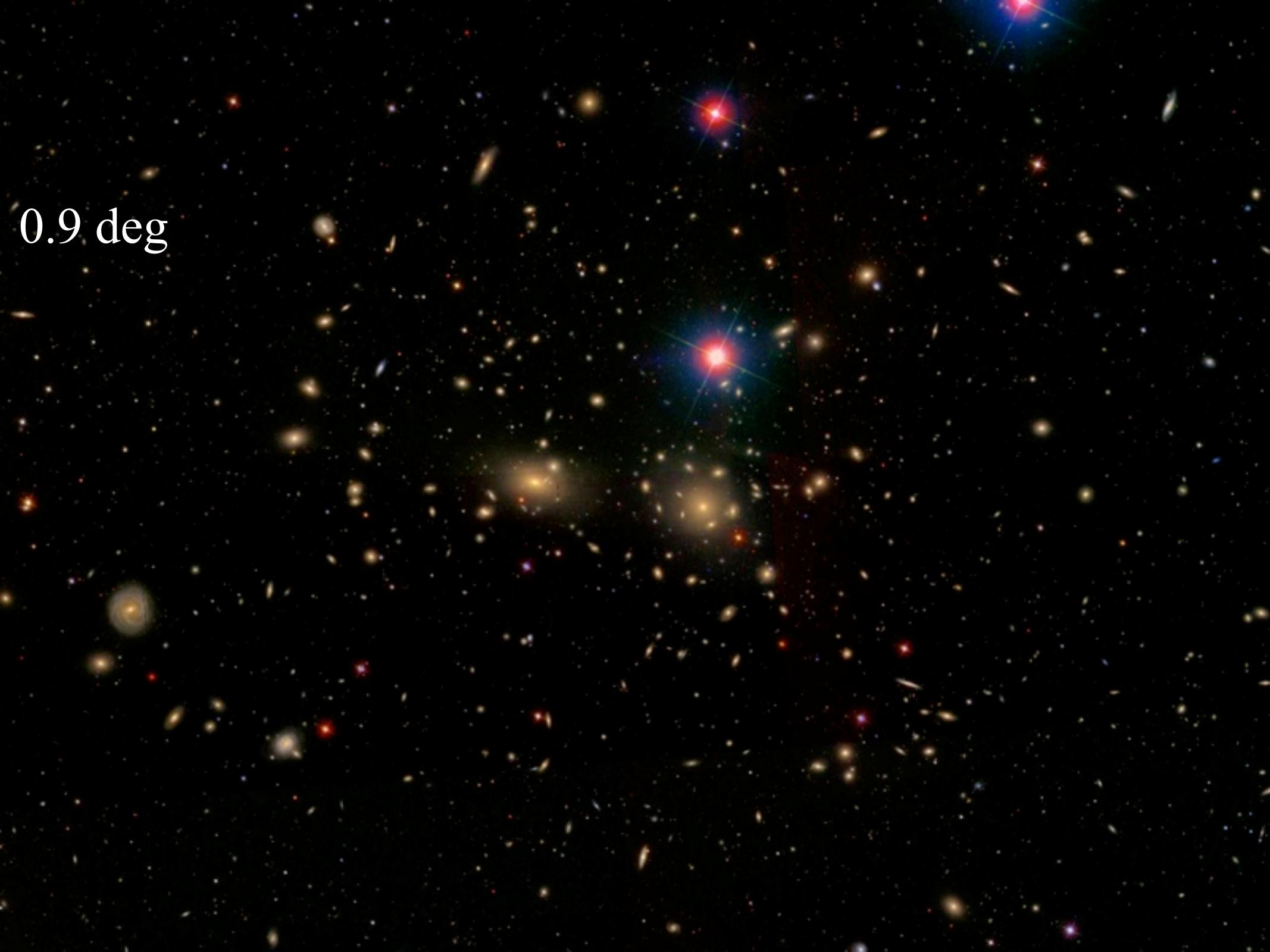
Spitzer Space Telescope • IRAC

ssc2006-02a

M31 = Andromeda: 3 deg
d = 780 kpc



0.9 deg

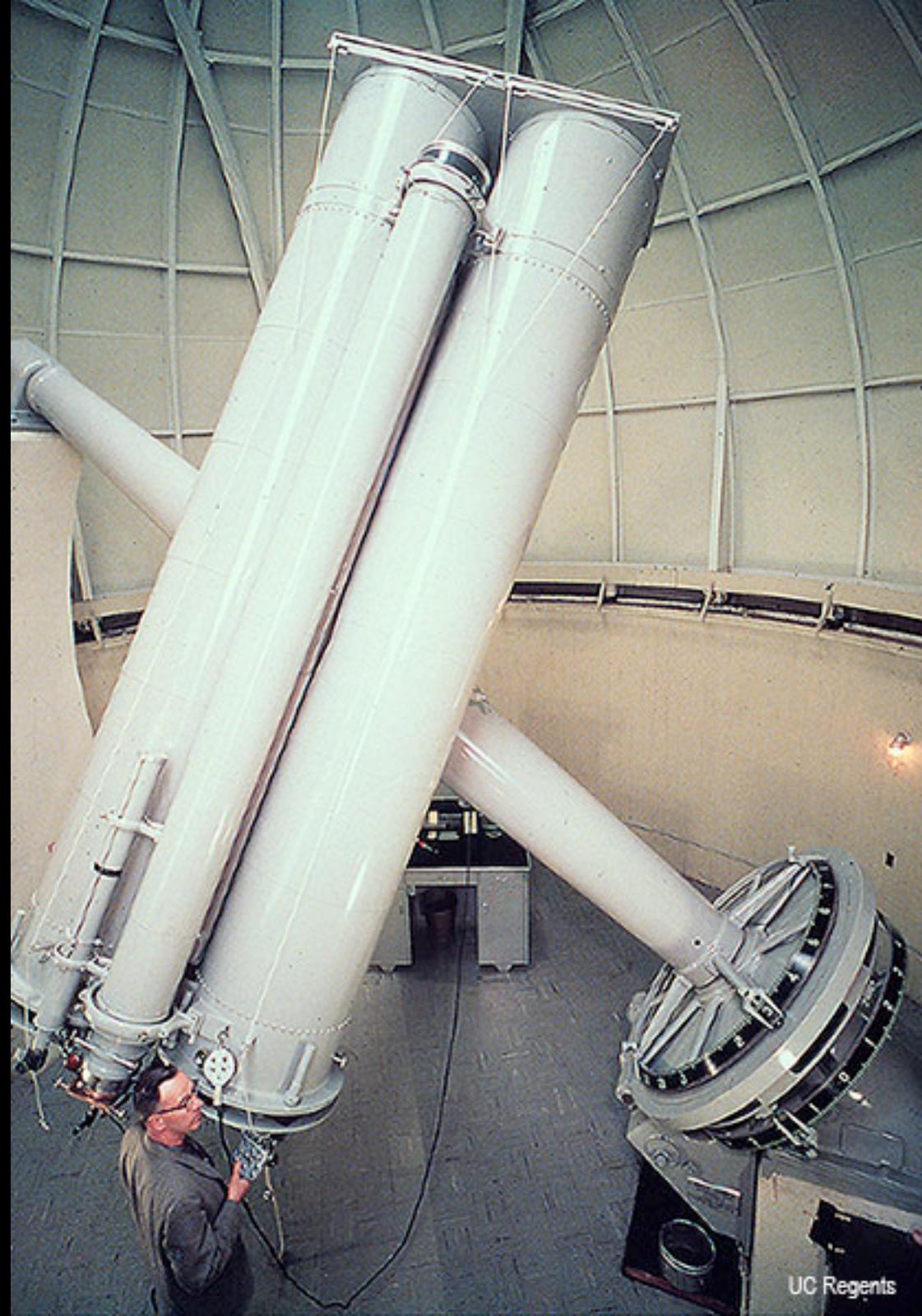


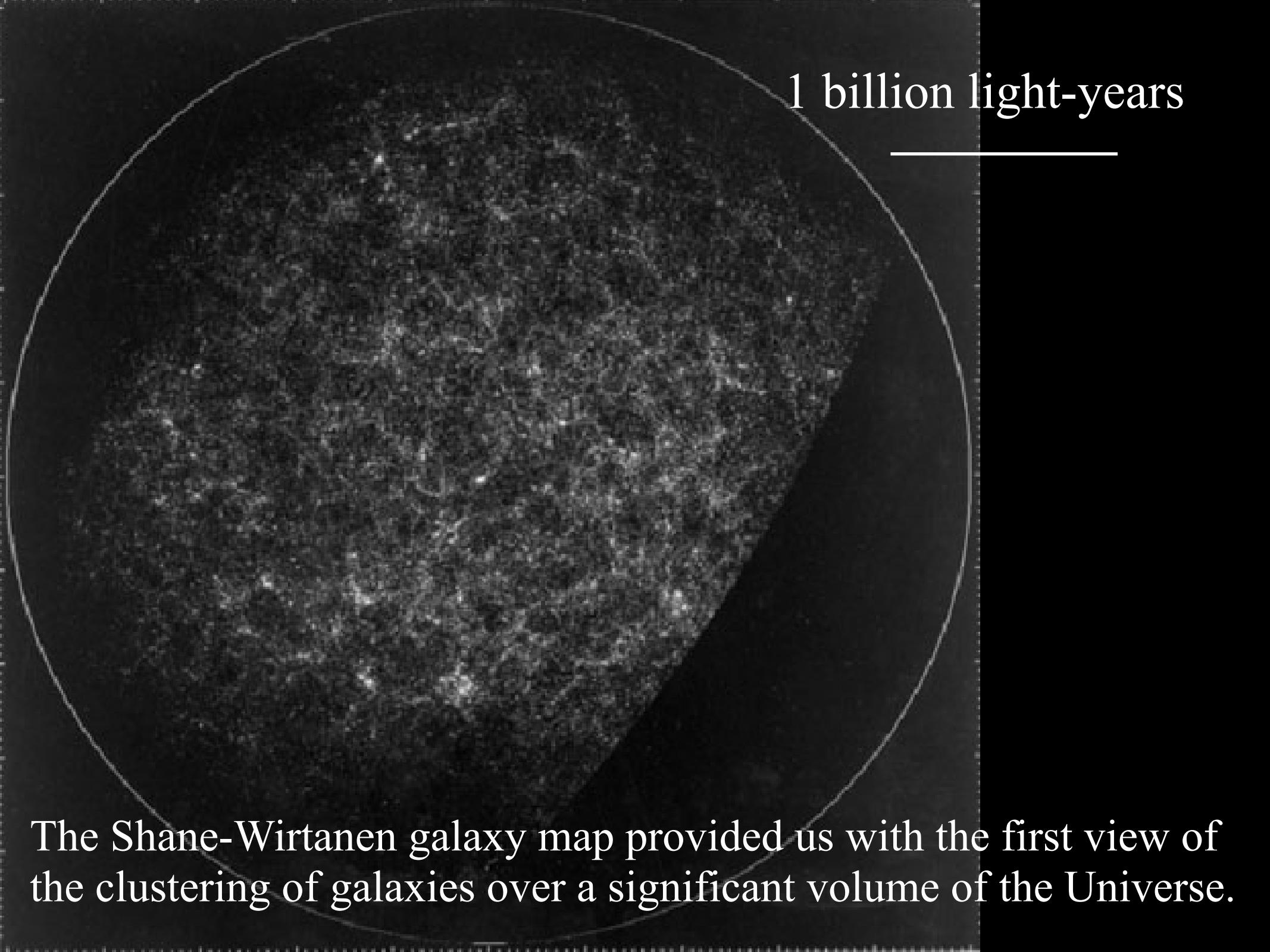
distances as light-travel times:

circumference of Earth	0.133 sec
distance to Moon	1.28 sec
circumference of Sun	15 sec
Sun	500 sec
Neptune	4 hours
diameter of Oort Cloud	10 months
α Centauri	4.3 years
center of Milky Way	27,000 years
Andromeda galaxy	2 million years
Virgo cluster	50 million years
Coma cluster	0.3 billion years
quasars	many billion years
cosmic horizon	13.7 billion years

This telescope at Lick Observatory made an atlas of the sky in the 1960's.

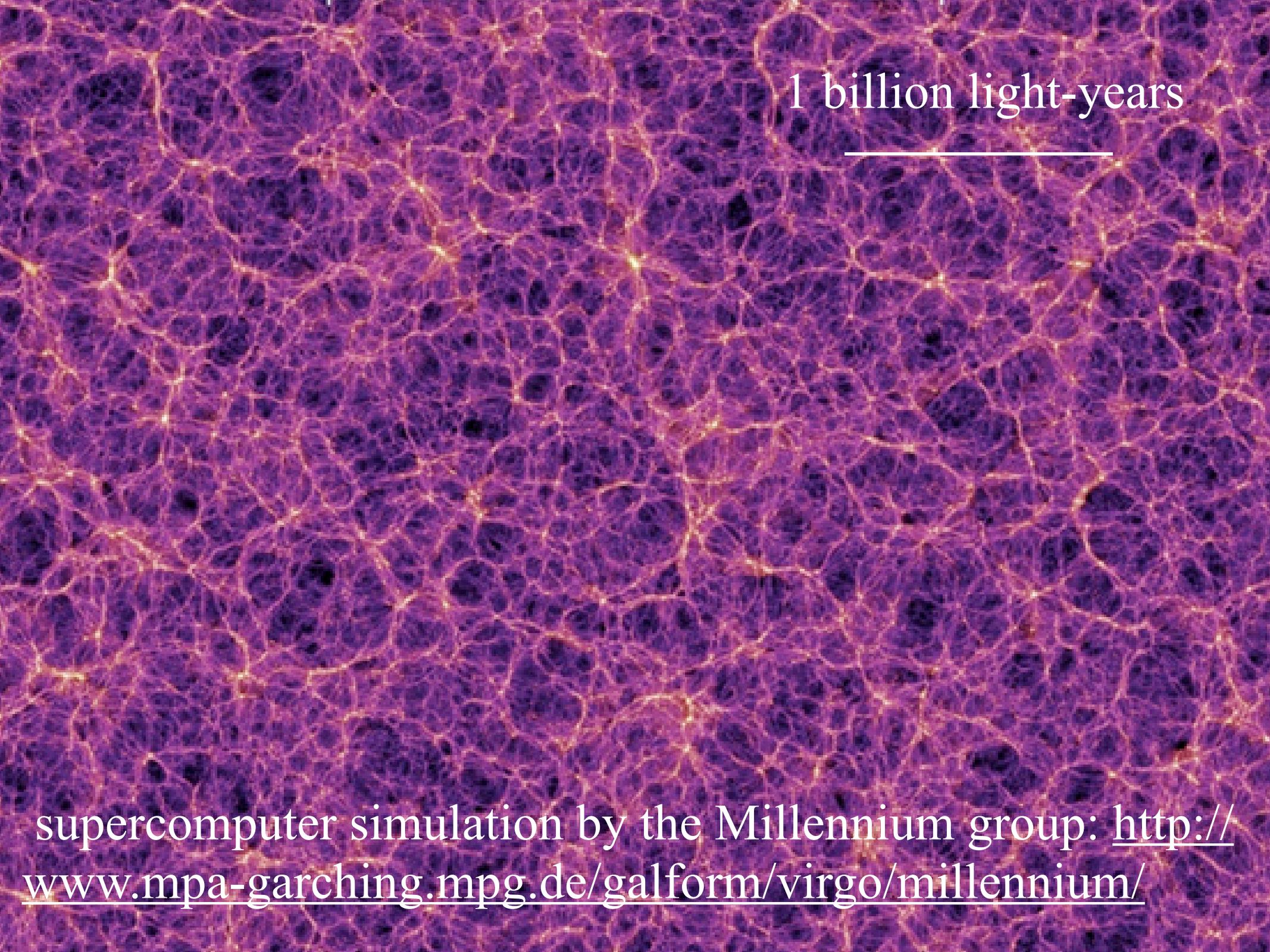
Carl Wirtanen and Donald Shane counted galaxies on the glass plates - one million galaxies in total.



A grayscale circular map showing the distribution of galaxies. The map is filled with a dense, granular texture of varying shades of gray, representing the density of galaxies across a volume. A thin white circle outlines the boundary of the map. In the top right corner, there is a vertical white line segment with a horizontal extension to its right, forming a scale bar.

1 billion light-years

The Shane-Wirtanen galaxy map provided us with the first view of the clustering of galaxies over a significant volume of the Universe.



1 billion light-years

supercomputer simulation by the Millennium group: [http://
www.mpa-garching.mpg.de/galform/virgo/millennium/](http://www.mpa-garching.mpg.de/galform/virgo/millennium/)

1932: brightest galaxies
(Shapley-Ames catalog)

$d \sim 10$ Mpc, $N \sim 1000$

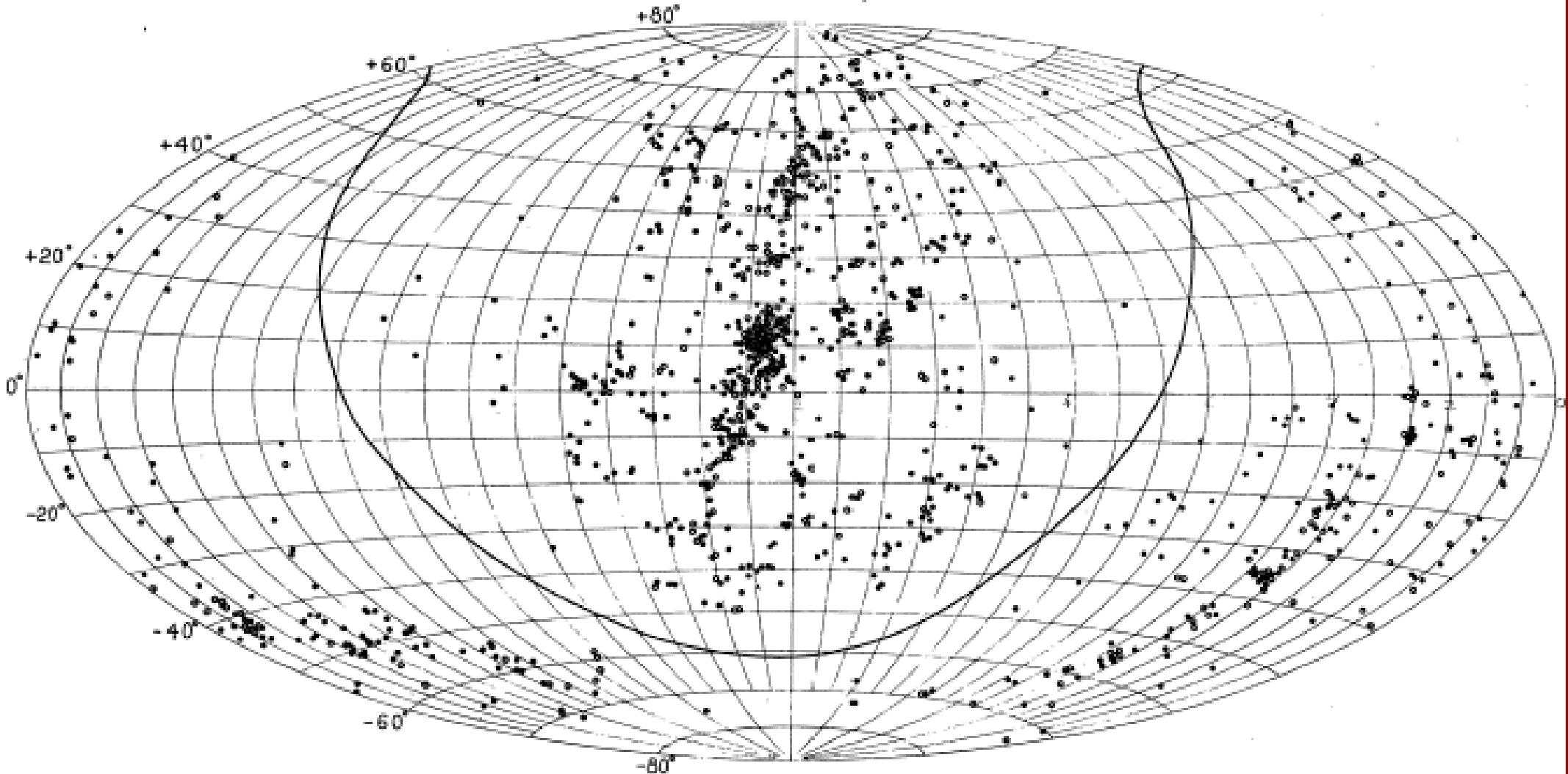


Figure 4